Nuclear Dynamics from Time Dependent Superfluid Local Density Approximation (TDSLDA)



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# **GOAL:** Description of nuclear dynamics far from equilibrium within the framework of TDDFT.

# Nuclear Skyrme functional

$$E = \int d^3 r \mathcal{H}(\mathbf{r})$$

where

$$\begin{aligned} \mathcal{H}(\mathbf{r}) &= C^{\rho}\rho^{2} + C^{s}\vec{s}\cdot\vec{s} + C^{\Delta\rho}\rho\nabla^{2}\rho + C^{\Delta s}\vec{s}\cdot\nabla^{2}\vec{s} + C^{\tau}(\rho\tau - \vec{j}\cdot\vec{j}) + \\ &+ C^{sT}(\vec{s}\cdot\vec{T} - \mathbf{J}^{2}) + C^{\nabla J}(\rho\vec{\nabla}\cdot\vec{J} + \vec{s}\cdot(\vec{\nabla}\times\vec{j})) + C^{\nabla s}(\vec{\nabla}\cdot\vec{s})^{2} + C^{\gamma}\rho^{\gamma} - \Delta\chi^{*} \end{aligned}$$

where

$$J_{i} = \sum_{k,l} \epsilon_{ikl} \mathbf{J}_{kl}$$
$$\mathbf{J}^{2} = \sum_{k,l} \mathbf{J}^{2}_{kl}$$

- density:  $\rho(\mathbf{r}) = \rho(\mathbf{r}, \mathbf{r'})|_{r=r'}$
- spin density:  $\vec{s}(\mathbf{r}) = \vec{s}(\mathbf{r}, \mathbf{r'})|_{r=r'}$
- current:  $\vec{j}(\mathbf{r}) = \frac{1}{2i}(\vec{\nabla} \vec{\nabla}')\rho(\mathbf{r}, \mathbf{r}')|_{r=r'}$
- spin current (2nd rank tensor):  $\mathbf{J}(\mathbf{r}) = \frac{1}{2i} (\vec{\nabla} \vec{\nabla}') \otimes \vec{s}(\mathbf{r}, \mathbf{r}')|_{r=r'}$
- kinetic energy density:  $\tau(\mathbf{r}) = \vec{\nabla} \cdot \vec{\nabla}' \rho(\mathbf{r}, \mathbf{r}')|_{r=r'}$
- spin kinetic energy density:  $\vec{T}(\mathbf{r}) = \vec{\nabla} \cdot \vec{\nabla}' \vec{s}(\mathbf{r}, \mathbf{r}')|_{r=r'}$
- anomalous (pairing) density:  $\chi(\mathbf{r}) = \chi(\mathbf{r}, \mathbf{r'})|_{r=r'}$

# **Treatment of the Coulomb potential**

$$7^{2}\Phi(\mathbf{r}) = 4\pi e^{2}\rho(\mathbf{r})$$
  
$$\Phi(\mathbf{r}) = \int d^{3}r' \frac{e^{2}\rho(\mathbf{r})}{|\mathbf{r} - \mathbf{r}'|}$$

$$\Phi(\mathbf{r}) = \int \frac{d^3k}{(2\pi)^3} \frac{e^2 \rho(\vec{k})}{k^2} \exp(i\vec{k} \cdot \mathbf{r}) = \frac{1}{27N_x N_y N_z} \sum_{\vec{k} \in L_x L_y L_z} e^2 \rho(\vec{k}) f(k) \exp(i\vec{k} \cdot \mathbf{r})$$

Defining an auxiliary potential f(r) on can get rid of spurious interaction with neighboring cells at the cost of performing FFT in 3 times larger box.

7



$$f(r) = 1/r \text{ for } r < \sqrt{L_x^2 + L_y^2 + L_z^2}$$
  
$$f(r) = 0 \text{ otherwise}$$

$$(r) = 0$$
 otherwise

$$f(k) = 4\pi \frac{1 - \cos(k\sqrt{L_x^2 + L_y^2 + L_z^2})}{k^2}$$

Castro, Rubio, Stott, arXive:0012024v1 However taking into account that FFT in a larger box means simply denser momentum space one can replace one FFT in 3 times larger box with 27 FFT's in the original box.

$$\begin{split} &\Phi(\mathbf{r}) = \\ &= \frac{1}{27N_x N_y N_z} \sum_{k,l,m=0}^2 \left[ \sum_{\vec{k} \in L^3} e^2 \rho_{klm}(\vec{k}) f\left(\vec{k} + \left(k\frac{2\pi}{3L_x}, l\frac{2\pi}{3L_y}, m\frac{2\pi}{3L_z}\right)\right) \exp(i\vec{k} \cdot \mathbf{r}) \right] \\ &\times \quad \exp\left(i \left(k\frac{2\pi}{3L_x}x + l\frac{2\pi}{3L_y}y + m\frac{2\pi}{3L_z}z\right)\right) \end{split}$$

## where

$$\rho_{klm}(\vec{k}) = \sum_{\mathbf{r} \in L^3} \rho(x, y, z) \exp\left(-i\left(k\frac{2\pi}{3L_x}x + l\frac{2\pi}{3L_y}y + m\frac{2\pi}{3L_z}z\right)\right) \exp(-i\vec{k} \cdot \mathbf{r})$$

#### Gain in computational cost:

$$27 \cdot N^3 Log N^3 < \left(3N\right)^3 Log \left(3N\right)^3$$

# Formalism for Time Dependent Phenomena: TDSLDA

Local density approximation (no memory terms – adiabatic TDDFT)

$$i\hbar\frac{\partial}{\partial t}\begin{pmatrix}u_{k\uparrow}(\mathbf{r},t)\\u_{k\downarrow}(\mathbf{r},t)\\v_{k\uparrow}(\mathbf{r},t)\\v_{k\downarrow}(\mathbf{r},t)\end{pmatrix} = \begin{pmatrix}h_{\uparrow,\uparrow}(\mathbf{r},t)&h_{\uparrow,\downarrow}(\mathbf{r},t)&0&\Delta(\mathbf{r},t)\\h_{\downarrow,\uparrow}(\mathbf{r},t)&h_{\downarrow,\downarrow}(\mathbf{r},t)&-\Delta(\mathbf{r},t)&0\\0&-\Delta^{*}(\mathbf{r},t)&-h_{\uparrow,\uparrow}^{*}(\mathbf{r},t)&-h_{\uparrow,\downarrow}^{*}(\mathbf{r},t)\end{pmatrix} \begin{pmatrix}u_{k\uparrow}(\mathbf{r},t)\\u_{k\downarrow}(\mathbf{r},t)\\v_{k\uparrow}(\mathbf{r},t)\\v_{k\downarrow}(\mathbf{r},t)\end{pmatrix}$$

Density functional contains normal densities, anomalous density (pairing) and currents:

$$E(t) = \int d^3r \left[ \varepsilon(n(\vec{r},t),\tau(\vec{r},t),\nu(\vec{r},t),\vec{j}(\vec{r},t)) + V_{ext}(\vec{r},t)n(\vec{r},t) + \dots \right]$$

- The system is placed on a large 3D spatial lattice.
- Derivatives are computed with FFTW
- Fully self-consistent treatment with fundamental symmetries respected (isospin, gauge, Galilean, rotation, translation)
- for TD high-accuracy and numerically stable Adams–Bashforth–Milne 5<sup>th</sup> order predictorcorrector-modifier algorithm with only 2 evaluations of the rhs per time step and with no matrix operations
- No symmetry restrictions
- Number of PDEs is of the order of the number of spatial lattice points
- Initial conditions for TDSLDA are generated from static SLDA code.
   In future: ground state may be generated through adiabatic switching and quantum friction (Bulgac et al. arXiv:1305.6891)

Single particle potential (Skyrme):

$$h(\mathbf{r}) = -\vec{\nabla} \cdot \left(B(\mathbf{r}) + \vec{\sigma} \cdot \vec{C}(\mathbf{r})\right) \vec{\nabla} + U(\mathbf{r}) + \frac{1}{2i} \left[\vec{W}(\mathbf{r}) \cdot (\vec{\nabla} \times \vec{\sigma}) + \vec{\nabla} \cdot (\vec{\sigma} \times \vec{W}(\mathbf{r}))\right] \\ + \vec{U}_{\sigma}(\mathbf{r}) \cdot \vec{\sigma} + \frac{1}{i} \left(\vec{\nabla} \cdot \vec{U}_{\Delta}(\mathbf{r}) + \vec{U}_{\Delta}(\mathbf{r}) \cdot \vec{\nabla}\right) \vec{\nabla}$$

where

$$\begin{split} B(\mathbf{r}) &= \frac{\hbar^2}{2m} + C^{\tau}\rho \\ \vec{C}(\mathbf{r}) &= C^{sT}\vec{s} \\ U(\mathbf{r}) &= 2C^{\rho}\rho + 2C^{\Delta\rho}\nabla^2\rho + C^{\tau}\tau + C^{\nabla J}\vec{\nabla}\cdot\vec{J} + C^{\gamma}(\gamma+2)\rho^{\gamma+1} \\ \vec{W}(\mathbf{r}) &= -C^{\nabla J}\vec{\nabla}\rho \\ \vec{U}_{\sigma}(\mathbf{r}) &= 2C^{s}\vec{s} + 2C^{\Delta s}\nabla^2\vec{s} + C^{sT}\vec{T} + C^{\nabla J}\vec{\nabla}\times\vec{j} \\ \vec{U}_{\Delta}(\mathbf{r}) &= C^{j}\vec{j} + \frac{1}{2}C^{\nabla j}\vec{\nabla}\times\vec{s} \end{split}$$

and pairing potential:

$$\Delta(\mathbf{r},t) = -g_{eff}(\mathbf{r})\chi(\mathbf{r},t)$$

Linear response regime: GDR of deformed nuclei

Box size: 32.5fm (mesh size: 1.25fm) Energy deposited into a nucleus: 45-50MeV Adiabatic switching of external perturbation: C\*exp[-(t-10)^2/2] Time window for Fourier transform: 1600 fm/c Time step: 0.12fm/c -> relative accuracy: 10^(-7) Photoabsorption cross section for heavy, deformed nuclei.

 $\begin{aligned} h_{\tau,\sigma\sigma}(\mathbf{r},t) &\Rightarrow h_{\tau,\sigma\sigma}(\mathbf{r},t) + F_{\tau}(\mathbf{r})f(t) & F_{\tau}(\mathbf{r}) = N_{\tau}\sin(\mathbf{k}\cdot\mathbf{r}_{\tau})/|\mathbf{k}| \\ S(E) &= \sum_{\nu} |\langle \nu|F|0\rangle|^2 \delta(E-E_{\nu}) \\ S(\omega) &= \operatorname{Im}\{\delta F(\omega)/[\pi f(\omega)]\} \\ \delta F(t) &= \langle \hat{F} \rangle_t - \langle \hat{F} \rangle_0 = \int d^3r \delta \rho(\mathbf{r},t) F(\mathbf{r}) f(t) = C \exp[-(t-10)^2/2] \end{aligned}$ 

### (gamma,n) reaction through the excitation of GDR

I.Stetcu, A.Bulgac, P. Magierski, K.J. Roche, Phys. Rev. C84 051309 (2011)



#### **Evolution of occupation probabilities**



FIG. 1. (Color online) The time-dependent proton and neutron occupation probabilities of a a number of quasiparticle states around the Fermi level for  $^{238}$ U calculated as described in the main text with SLy4.

Occupation probabilities vary significantly in time.

Pairing has to be treated fully selfconsistently!

Beyond linear regime: Relativistic Coulomb excitation **Coupling to e.m. field:** 

$$\vec{E} = -\vec{\nabla}\phi - \frac{1}{c}\frac{\partial\vec{A}}{\partial t}$$
$$\vec{B} = \vec{\nabla}\times\vec{A}$$
$$\vec{\nabla}\psi \rightarrow \vec{\nabla}_A\psi = \left(\vec{\nabla} - i\frac{e}{\hbar c}\vec{A}\right)\psi$$
$$\vec{\nabla}\psi^* \rightarrow \vec{\nabla}_{-A}\psi^* = \left(\vec{\nabla} + i\frac{e}{\hbar c}\vec{A}\right)\psi^*$$
$$i\hbar\frac{\partial}{\partial t}\psi \rightarrow \left(i\hbar\frac{\partial}{\partial t} - e\phi\right)\psi$$

which implies that  $\vec{\nabla}\psi\psi^* \to \vec{\nabla}\psi\psi^*$ .

Consequently the densities change according to:

- density:  $\rho_A(\mathbf{r}) = \rho_A(\mathbf{r})$
- spin density:  $\vec{s}_A(\mathbf{r}) = \vec{s}(\mathbf{r})$
- current:  $\vec{j}_A(\mathbf{r}) = \vec{j}(\mathbf{r}) \frac{e}{\hbar c} \vec{A} \rho(\mathbf{r})$
- spin current (2nd rank tensor):  $\mathbf{J}_A(\mathbf{r}) = \mathbf{J}(\mathbf{r}) \frac{e}{\hbar c} \vec{A} \otimes \vec{s}(\mathbf{r})$
- spin current (vector):  $\vec{J}_A(\mathbf{r}) = \vec{J}(\mathbf{r}) \frac{e}{\hbar c} \vec{A} \times \vec{s}(\mathbf{r})$

• kinetic energy density: 
$$\tau_A(\mathbf{r}) = \left(\vec{\nabla} - i\frac{e}{\hbar c}\vec{A}\right) \cdot \left(\vec{\nabla}' + i\frac{e}{\hbar c}\vec{A}\right)\rho(\mathbf{r},\mathbf{r}')|_{r=r'}$$
  
=  $\tau(\mathbf{r}) - 2\frac{e}{\hbar c}\vec{A}\cdot\vec{j}(\mathbf{r}) + \frac{e^2}{\hbar^2 c^2}|\vec{A}|^2\rho(\mathbf{r}) = \tau(\mathbf{r}) - 2\frac{e}{\hbar c}\vec{A}\cdot\vec{j}_A(\mathbf{r}) - \frac{e^2}{\hbar^2 c^2}|\vec{A}|^2\rho(\mathbf{r})$ 

• spin kinetic energy density:  $\vec{T}_A(\mathbf{r}) = \left(\vec{\nabla} - i\frac{e}{\hbar c}\vec{A}\right) \cdot \left(\vec{\nabla}' + i\frac{e}{\hbar c}\vec{A}\right)\vec{s}(\mathbf{r},\mathbf{r}')|_{r=r'}$ =  $\vec{T}(\mathbf{r}) - 2\frac{e}{\hbar c}\vec{A}^T \cdot \mathbf{J}(\mathbf{r}) + \frac{e^2}{\hbar^2 c^2}|\vec{A}|^2\vec{s}(\mathbf{r}) = \vec{T}(\mathbf{r}) - 2\frac{e}{\hbar c}\vec{A}^T \cdot \mathbf{J}_A(\mathbf{r}) - \frac{e^2}{\hbar^2 c^2}|\vec{A}|^2\vec{s}(\mathbf{r})$ 

## **Relativistic Coulomb excitation**

# $^{238}U + ^{238}U \rightarrow ^{238}U * +...$ at about 700 MeV/n



The coordinate transformation has been applied to keep CM in the center of the box at all times.

**Energy deposited for two nuclear orientations (y – perpendicular, z – parallel)** 

#### Impact parameter b=12.2fm



Impact parameter b=16.2fm



### Neutron emission

#### Impact parameter b=12.2fm

146

Neutron number

Number of neutrons in two shells surrounding nucleus for two nuclear orientations with respect to incoming projectile:





# Protons

# **Internal nuclear excitations**

### Electric dipole moment (along two axes: y, z) as a function of time



Oscillations are damped due to the one-body dissipation mechanism

# **Electromagnetic radiation from excited nucleus**

$$\begin{split} \rho(\mathbf{r},t) &= \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \rho(\mathbf{r},\omega) \exp(-i\omega t) \\ \vec{j}(\mathbf{r},t) &= \frac{1}{2\pi} \int_{-\infty}^{\infty} d\omega \vec{j}(\mathbf{r},\omega) \exp(-i\omega t) \\ \end{bmatrix} \mathbf{From TDSLDA} \\ \vec{B}(\mathbf{r},\omega) &= \frac{ie}{c} \frac{\exp(ikr)}{r} \int d^3r' \vec{k} \times \vec{j}(\mathbf{r}',\omega) \exp(-i\vec{k}\cdot\mathbf{r}') = \frac{ie}{c} \frac{\exp(ikr)}{r} \vec{k} \times \vec{j}(\vec{k},\omega) \\ \vec{E}(\mathbf{r},\omega) &= \frac{ie}{c} \frac{\exp(ikr)}{r} \frac{\mathbf{r}}{r} \times \int d^3r' (\vec{j}(\mathbf{r}',\omega) \times \vec{k}) \exp(-i\vec{k}\cdot\mathbf{r}') = \frac{ie}{c} \frac{\exp(ikr)}{r} \frac{\mathbf{r}}{r} \times \left(\vec{j}(\vec{k},\omega) \times \vec{k}\right) \\ \frac{dP}{d\Omega}(t) &= \frac{e^2}{4\pi c} \left| \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} (\vec{k} \times \vec{j}(\vec{k},\omega)) \exp(-i\omega(t-r/c)) \right|^2 \\ \mathbf{Angular distribution of radiated power} \end{split}$$

$$\frac{dE}{d\Omega d\omega}(\omega) = \frac{e^2}{4\pi^2 c} \left| \vec{k} \times \vec{j}(\vec{k},\omega) \right|^2 = \frac{e^2}{4\pi^2 c} \left| \int d^3 r \left( \nabla \times \vec{j}(\mathbf{r},\omega) \right) \exp(-i\vec{k} \cdot \mathbf{r}) \right|^2$$
Angular distribution and ferquency distribution of emitted radiation

#### In practice it is better to perform multipole expansion:

$$\frac{dE}{d\omega} = \frac{4e^2}{c} \sum_{l,m} |\vec{b}_{lm}(k,\omega)|^2$$

$$P(t+r/c) = \int \frac{dP}{d\Omega} (t+r/c) d\Omega = \frac{e^2}{\pi c} \sum_{l,m} \left| \int_{-\infty}^{\infty} \vec{b}_{lm}(k,\omega) \exp(-i\omega t) d\omega \right|^2$$

$$\vec{b}_{lm}(k,t) = \int d^3 r \vec{b}(\mathbf{r},t) j_l(kr) Y_{lm}^*(\hat{r})$$

$$\vec{b}_{lm}(k,\omega) = \int_{-\infty}^{\infty} \vec{b}_{lm}(k,t) \exp(i\omega t) dt$$

#### Electromagnetic radiation rate due to the internal motion



# Electromagnetic radiation due to the internal nuclear motion



# **Summary**

- TDSLDA is a flexible tool to study nuclear dynamics.
- Pairing field is treated on the same footing like single particle potentais (no frozen occupation number approximation).
- Nuclear excitation modes (beyond linear response!) can be identified from e.m. radiation.
- Various nonequilibrium nuclear processes can be studied:
  - Nuclear large amplitude collective motion (LACM
  - (induced) nuclear fission
  - Excitation of nuclei with gamma rays and neutrons
  - Coulomb excitation of nuclei with relativistic heavy-ions
  - Nuclear reactions, fusion between colliding heavy-ions
  - Neutron star crust and dynamics of vortices and their pinning mechanism

# **Current capabilities of the code:**

- volumes of the order of (L = 80<sup>3</sup>) capable of simulating time evolution of 42000 neutrons at saturation density (possible application: neutron stars)
- capable of simulating up to times of the order of 10<sup>-19</sup> s (a few million time steps)

<u>CPU vs GPU on Titan ≈ 15 speed-up</u> (likely an additional factor of 4 possible)
 Eg. for 137062 two component wave functions:
 CPU version (4096 nodes x 16 PEs) - 27.90 sec for 10 time steps
 GPU version (4096 PEs + 4096GPU) - 1.84 sec for 10 time steps