Towards exascale simulations of quantum superfluids – new prospects for modelling nuclear processes







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GOAL:

Description of <u>superfluid dynamics</u> of fermionic systems <u>far from equilibrium</u> based on microscopic theoretical framework.

Microscopic framework = explicit treatment of fermionic degrees of freedom.

Why Time Dependent Density Functional Theory (TDDFT)?

We need to describe the time evolution of (externally perturbed) spatially inhomogeneous, superfluid Fermi system.

Within current computational capabilities TDDFT allows to describe real time dynamics of strongly interacting, superfluid systems of <u>hundred of thousands</u> fermions.

Main Theoretical Tool

THEOREM: There exist an universal density functional of particle density.





2012

DFT has been developed and used mainly to describe normal (non-superfluid) electron systems – more than <u>50 years old theory</u>:

DFT - Kohn and Hohenberg, 1964 LDA - Kohn and Sham, 1965 Extension for superconductors/superfluids: DFT - Oliveira, Gross, Kohn (1988), TDDFT - Wacker, Kummel, Gross (1994) SLDA, TDSLDA - Bulgac (2002)

TDDFT equations with local pairing field (TDSLDA):

$$i\hbar\frac{\partial}{\partial t} \begin{pmatrix} u_{k\uparrow}(\mathbf{r},t) \\ u_{k\downarrow}(\mathbf{r},t) \\ v_{k\uparrow}(\mathbf{r},t) \\ v_{k\downarrow}(\mathbf{r},t) \end{pmatrix} = \begin{pmatrix} h_{\uparrow,\uparrow}(\mathbf{r},t) & h_{\uparrow,\downarrow}(\mathbf{r},t) & 0 & \Delta(\mathbf{r},t) \\ h_{\downarrow,\uparrow}(\mathbf{r},t) & h_{\downarrow,\downarrow}(\mathbf{r},t) & -\Delta(\mathbf{r},t) & 0 \\ 0 & -\Delta^*(\mathbf{r},t) & -h^*_{\uparrow,\uparrow}(\mathbf{r},t) & -h^*_{\uparrow,\downarrow}(\mathbf{r},t) \\ \Delta^*(\mathbf{r},t) & 0 & -h^*_{\uparrow,\downarrow}(\mathbf{r},t) & -h^*_{\downarrow,\downarrow}(\mathbf{r},t) \end{pmatrix} \begin{pmatrix} u_{k\uparrow}(\mathbf{r},t) & u_{k\downarrow}(\mathbf{r},t) \\ u_{k\downarrow}(\mathbf{r},t) \\ v_{k\uparrow}(\mathbf{r},t) \\ v_{k\downarrow}(\mathbf{r},t) \end{pmatrix}$$

The form of h(r,t) and $\Delta(r,t)$ is determined by EDF (Energy Density Functional)

- •The system is placed on a large 3D spatial lattice.
- No symmetry restrictions.
- Number of PDEs is of the order of the number of spatial lattice points.

Table 1: Comparison of profit gained by using GPUs instead of CPUs for two example lattices. The timing was obtained on Titan supercomputer. Note, Titan has 16x more CPUs than GPUs.

	CPU impleme	ntation	GPU impleme	ntation	
Number of H	FB				
N _x N _y N _z equatio	ns # of CPUs	time per step	# of GPUs	time per step	SPEEDUP
48 ³ 110,5	592 110,592	3.9 sec	6,912	0.39 sec	10
64 ³ 262,1	.44 262,144	20 sec	16,384	0.80 sec	25

The main advantage of TDSLDA over TDHF (+TDBCS) is related to the fact that in TDSLDA the pairing correlations are described as a true <u>complex field which has its own modes of excitations</u>, which include spatial variations of both amplitude and phase. Therefore in TDSLDA description the evolution of nucleon Cooper pairs is treated consistently with other one-body degrees of freedom.

Selected supercomputers (CPU+GPU) currently in use:



Piz Daint: 25.3 PFlops (Swiss National Supercomputing Centre)

HA-PACS: 0.802 PFlops (University of Tsukuba)





Tsubame: 5.7 PFlops (Tokyo Institute of Technology)

TSUBAME

Titan: 27 PFlops (ORNL Oak Ridge)



Advancing the Era of Accelerated Computing





Areas of applications



Ultracold atomic (fermionic) gases. Unitary regime. Dynamics of quantum vortices, solitonic excitations, quantum turbulence.





 $\frac{\Delta}{-\!\!-\!\!-} \leq 0.1 \!-\! 0.2$ \mathcal{E}_{F}

Astrophysical applications. Modelling of neutron star interior (glitches): vortex dynamics, dynamics of inhomogeneous nuclear matter (in strong magnetic fields).





Nuclear physics. Induced nuclear fission, fusion, collisions.





Examples of applications:

- Nuclear induced fission
- Collisions of medium or heavy superfluid nuclei
- Dynamics of inhomogeneous nuclear matter in neutron stars
- Dynamics of topological excitations in ultracold fermionic gases (both unpolarized and spin-imbalanced)

Fission dynamics of ²⁴⁰Pu

Initial configuration of ${}^{240}Pu$ is prepared beyond the barrier at quadrupole deformation Q=165b and excitation energy E=8.08 MeV: Accelerations in quadrupole and octu-



Accelerations in quadrupole and octupole moments along the fission path



Note that despite the fact that nucleus is already <u>beyond the saddle point</u> the collective motion on the time scale of 1000 fm/c and larger is characterized by <u>the constant velocity</u> (*see red dashed line for an average acceleration*) till the very last moment before splitting. On times scales, of the order of 300 fm/c and shorter, the collective motion is a subject to random-like kicks indicating strong coupling to internal d.o.f



Induced fission of 240Pu

The lighter fragment is more excited (and strongly deformed) than the heavier one.

Energies are not shared proportionally to mass numbers of the fragments!

TKE = 177.80	$-0.3489E_n$	[in MeV],
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Nuclear data evaluation, Madland (2006)

Calculated total kinetic energies (TKE) of the fragments slightly underestimate the observed values by no more than: 1 - 3 MeV !

J. Grineviciute, et al. (in preparation) see also:

A. Bulgac, P. Magierski, K.J. Roche, and I. Stetcu, Phys. Rev. Lett. 116, 122504 (2016)

Nuclear collisions

Collisions of superfluid nuclei having <u>different phases</u> of the <u>pairing fields</u>

Inspired by experiments on ultracold atomic gases: merging two 6Li clouds





Creation of a "<u>heavy soliton</u>" after merging two superfluid atomic clouds. T. Yefsah et al., Nature 499, 426 (2013).

And recently detailed analysis of **solitonic cascade** has been performed in experiment at MIT: M.J.H. Ku et al. Phys. Rev. Lett. 116, 045304 (2016)

Wlazłowski, Sekizawa, Magierski (in preparation)

<u>In the context of nuclear systems the main questions are:</u> -how a possible solitonic structure can be manifested in nuclear system? -what observable effect it may have on heavy ion reaction: kinetic energies of fragments, capture cross section, etc.?

Clearly, we cannot control phases of the pairing field in nuclear experiments and the possible signal need to be extracted after averaging over the phase difference.



Y. Hashimoto, G. Scamps, Phys. Rev. C94, 014610(2016) – TDHFB studies of small systems: 200+200 reaction produced negligible effect.

Estimates for the magnitude of the effect

At first one may think that the magnitude of the effect is determined by the nuclear pairing energy which is of the order of MeV's in atomic nuclei (according to the expression):

$$\frac{1}{2}g(\varepsilon_F)|\Delta|^2$$
; $g(\varepsilon_F)$ - density of states

On the other hand the energy stored in the junction can be estimated from Ginzburg-Landau (G-L) approach:

For typical values characteristic for two heavy nuclei: $E_i \approx 30 MeV$



Creation of <u>the solitonic structure</u> between colliding nuclei prevents energy transfer to internal degrees of freedom and consequently <u>enhances</u> the kinetic energy of outgoing fragments. Surprisingly, the <u>gauge angle dependence</u> from the G-L approach is perfectly well reproduced in <u>the kinetic energies of outgoing fragments</u>!



Modelling neutron star interior - towards microscopic foundations of the neutron star crust dynamics

<u>GOAL</u>: Construct large scale model of neutron star interior (in particular <u>neutron star crust</u>), based on microscopic input from nuclear theory.

MICROSCOPIC INPUTS NEEDED:

- vortex-impurity interaction,
- effective masses of nuclear impurities,
- couplings between lattice vibrations and neutron superfluid medium,



Glitch: a sudden increase of the rotational frequency

Glitches in the Vela pulsar



V.B. Bhatia, A Textbook of Astronomy and Astrophysics with Elements of Cosmology, Alpha Science, 2001.

glitch phenomenon=a sudden speed up of rotation. To date more than 300 glitches have been detected in more than 100 pulsars

Glitch phenomenon is commonly believed to be related to rearrangement of vortices in the interior of neutron stars (Anderson, Itoh, Nature 256, 25 (1975)) It would require however a correlated behavior of huge number of quantum vortices and the mechanism of such collective rearrangement is still a mystery.

Vortex – impurity interaction

Static approach



Pinning energy is obtained as a result of substraction of two large numbers!

Dynamic approach

The external potential keeps the nucleus moving along the straight line with a constant velocity below the critical velocity.



G. Wlazłowski, K. Sekizawa, P. Magierski, A. Bulgac, M.M. Forbes, Phys. Rev. Lett. 117, 232701(2016)

Dynamics of ultracold atomic (fermionic) gases

✓In 1999 DeMarco and Jin created a degenerate atomic Fermi gas.

 ✓ In 2005 Zwierlein/Ketterle group observed quantum vortices which survived when passing from BEC to unitarity – evidence for superfluidity!

system of fermionic ⁶Li atoms

Feshbach resonance: B=834G





Figure 2 | Vortices in a strongly practice of the remionic atoms on the BEC- and the BCS-side of the Feshbach resonance. At the given field, the cloud of lithium atoms was stirred for 300 ms (a) or 500 ms (b-h) followed by an equilibration time of 500 ms. After 2 ms of ballistic expansion, the

magnetic field was ramped to 735 G for imaging (s magnetic fields were 740 G (a), 766 G (b), 792 G (c 843 G (f), 853 G (g) and 863 G (h). The field of view 880 μ m × 880 μ m.

M.W. Zwierlein *et al.*, Nature, 435, 1047 (2005)

What is a unitary gas?

A gas of interacting fermions is in the unitary regime if the average separation between particles is large compared to their size (range of interaction), but small compared to their scattering length.

$$n r_0^3 << 1 \quad n |a|^3 >> 1$$

$$n - particle density$$

$$a - scattering length$$

$$r_0 - effective range$$

$$n - particle density$$

$$r_0 - effective range$$

$$n - particle density$$

$$r_0 - effective range$$

$$REGIME$$
System is dilute but

strongly interacting!

Universality:
$$E(x) = \xi(x)E_{FG}$$
; $x = \frac{T}{\varepsilon_F}$
 $\xi(0) = 0.37(1)$ - Exp. estimate
 E_{FG} - Energy of noninteracting Fermi gas

Stirring the atomic cloud with stirring velocity lower than the critical velocity





Bulgac, Luo, Magierski, Roche, Yu, Science 332, 1288 (2011)

Stirring the atomic cloud with stirring velocity exceeding the critical velocity





Bulgac, Luo, Magierski, Roche, Yu, Science 332, 1288 (2011)



Moreover with TDDFT we can reproduce the sequence of topological excitations observed experimentally (M.H.J. Ku et al. Phys. Rev. Lett. 113, 065301 (2014)).

Wlazłowski, et al., Phys. Rev. A91, 031602 (2015)

Summarizing

 TDDFT extended to superfluid systems and based on the local densities offers a flexible tool to study quantum superfluids far from equilibrium.

Future plans:

- Ultracold atoms: investigation of <u>quantum turbulence</u> in Fermi systems; topological excitations in <u>spin-polarized</u> atomic gases in the presence of <u>LOFF phase</u>.
 Neutron star: Provide a link between <u>large scale models</u> of neutron stars and microscopic studies; towards the first simulation of the glitch phenomenon based on microscopic input.
- Nuclear physics: induced fission and fusion processes <u>from more</u> <u>phenomenology and adjusted parameters to more</u> <u>fundamental theory and increased predictive power</u>; search for new effects related to <u>pairing dynamics</u> in <u>nuclear nuclear processes</u>. Extension of TDDFT (account for dissipation and fluctuations of one-body observables):
 stochastic extension of TDDFT
 - Baranger-Veneroni prescription